

# VUV & SX Optics

Kenta Amemiya

Photon Factory, High Energy Accelerator Research Organization

## Outline

### 1. Mirrors

1.1. focusing & collimation

1.2. substrate materials

1.3. coating & reflectivity

### 2. Diffraction Gratings

2.1. principle - wavelength dispersion -

2.2. energy resolution

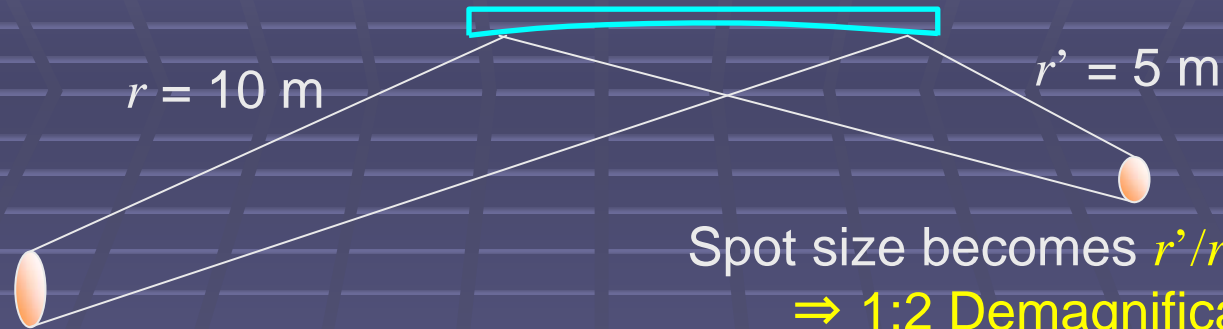
- dispersion and focus -

2.3. fabrication of gratings

## 1.1. focusing & collimation

### Demagnification & Aberration

$$R = \frac{2}{(1/r + 1/r') \cos \theta}$$



Spot size becomes  $r'/r = 1/2$   
 $\Rightarrow$  1:2 Demagnification

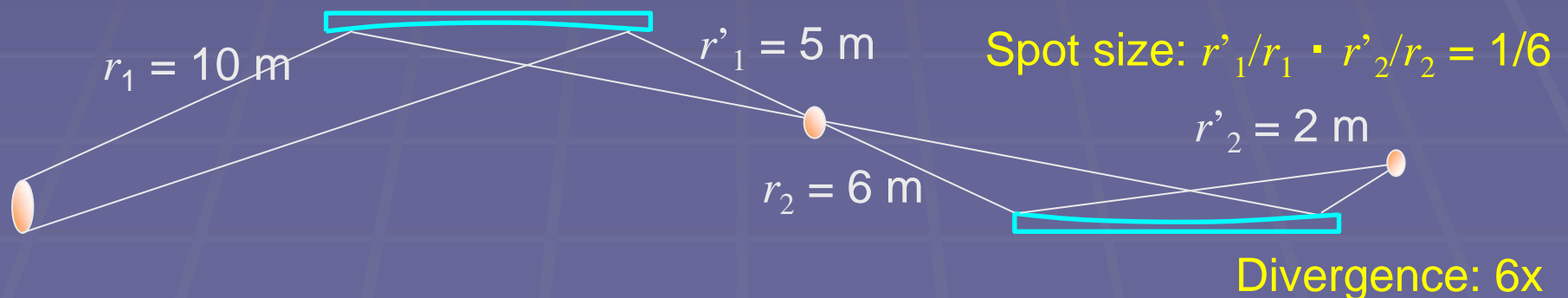
\* Divergence becomes 2 times larger

Elliptical mirror  $\Rightarrow$  perfect focus (determined by demagnification only)

Use of a cylindrical or spherical mirror  $\Rightarrow$  poor focus (aberration)

Larger divergence (larger illumination area)  $\Rightarrow$  larger aberration

### Combination of focusing mirrors

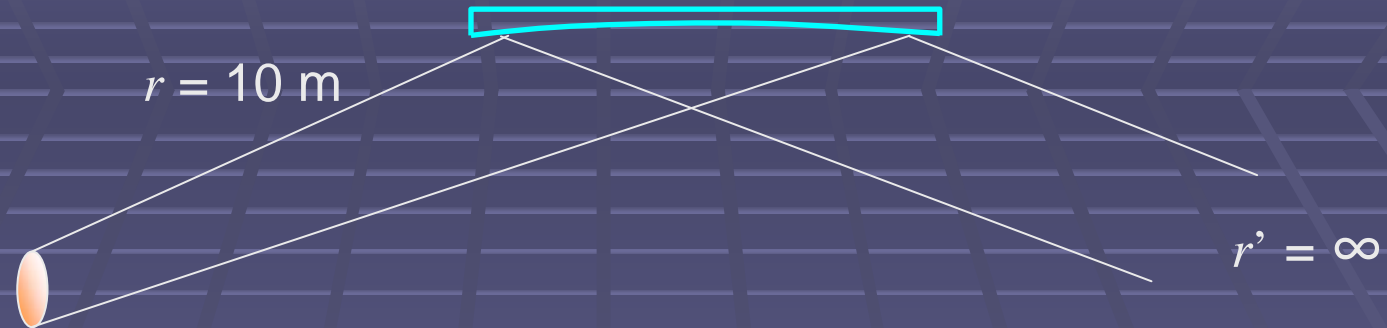


Spot size:  $r'_1/r_1 \cdot r'_2/r_2 = 1/6$

Divergence: 6x

## 1.1. focusing & collimation

### Collimation

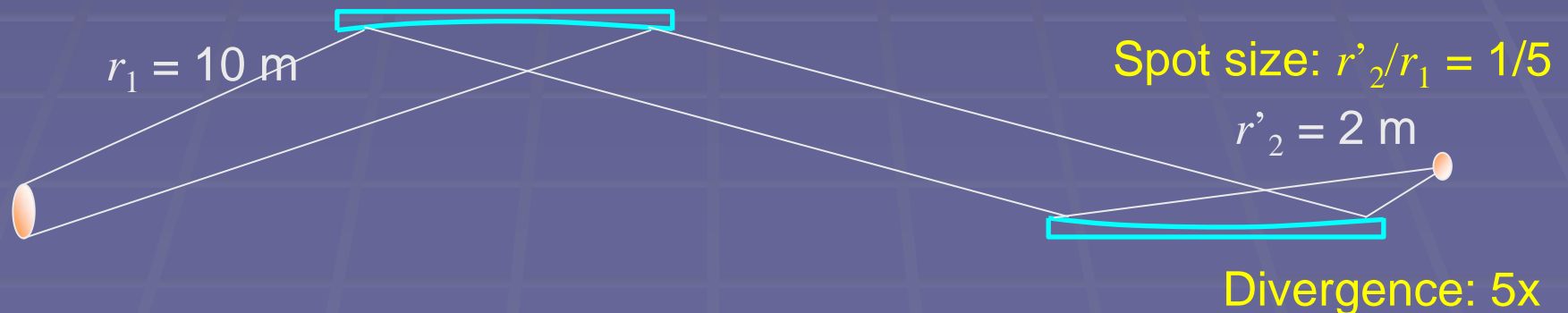


**Parabolic mirror**  $\Rightarrow$  perfect collimation (spot-size limited)

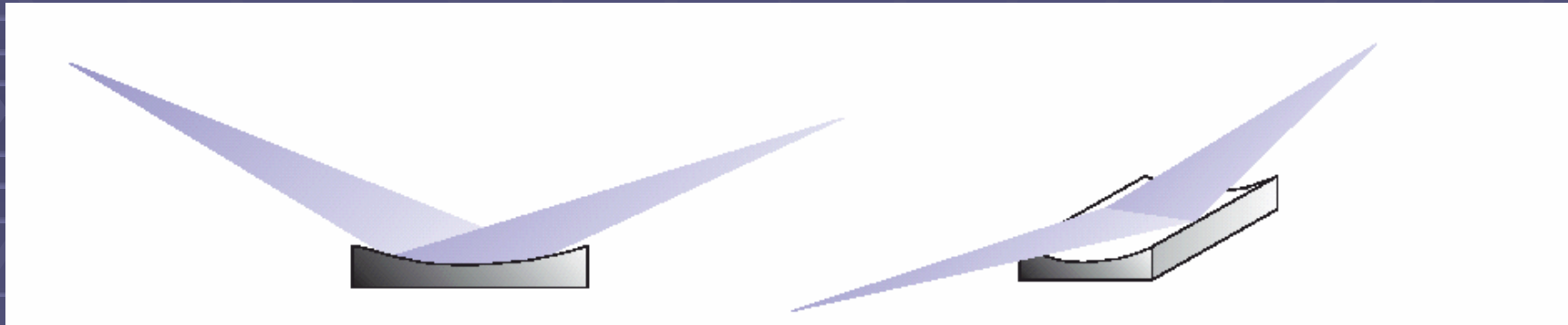
Use of a cylindrical or spherical mirror  $\Rightarrow$  poor collimation (**aberration**)

Larger divergence (larger illumination area)  $\Rightarrow$  larger aberration

### Combination of collimating mirrors



## 1.1. focusing & collimation



Tangential focusing

Sagittal focusing

**Aberration:** smaller in **tangential** focusing

\* Sagittal focusing is often adopted undulator beamlines

**Radius:** smaller in **sagittal** focusing

e.g. 88 deg,  $r = 10$  m,  $r' = 5$  m

$$R = \frac{2}{(1/r + 1/r') \cos \theta}$$

$$\rho = \frac{2 \cos \theta}{(1/r + 1/r')}$$

⇒  $R = 190$  m (tangential focusing),  $\rho = 0.23$  m (sagittal focusing)

**Higher precision** in **sagittal** focusing

e.g.  $R = 190$  m  $\pm$  3%,  $\rho = 0.23$  m  $\pm$  < 0.5 %

**Slope error effects:** smaller in **sagittal** focusing

## 1.2. substrate materials

### Needs:

Easy to fabricate: precise control of the mirror shape

Low defects, pores

Hardness: small distortion

High thermal conductivity: cooling efficiency

Low thermal expansion: against a heat load

### Typical materials

**Si**: for high heat load, with cooling

**SiO<sub>2</sub>**: without cooling

\* SiO<sub>2</sub> is suitable for mirror current measurements

## 1.3. coating & reflectivity

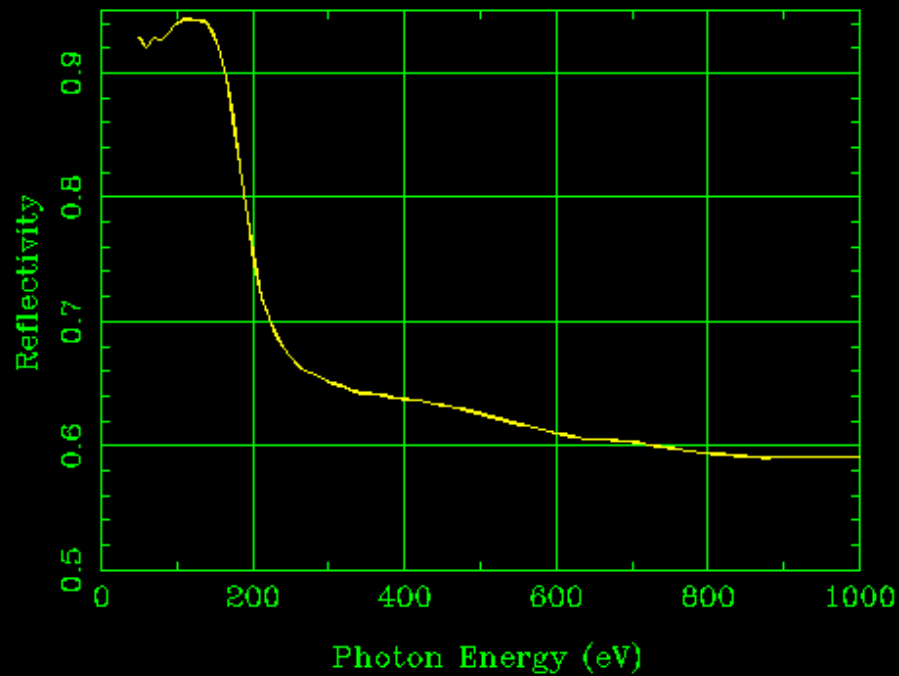
Coating layer:

High reflectivity

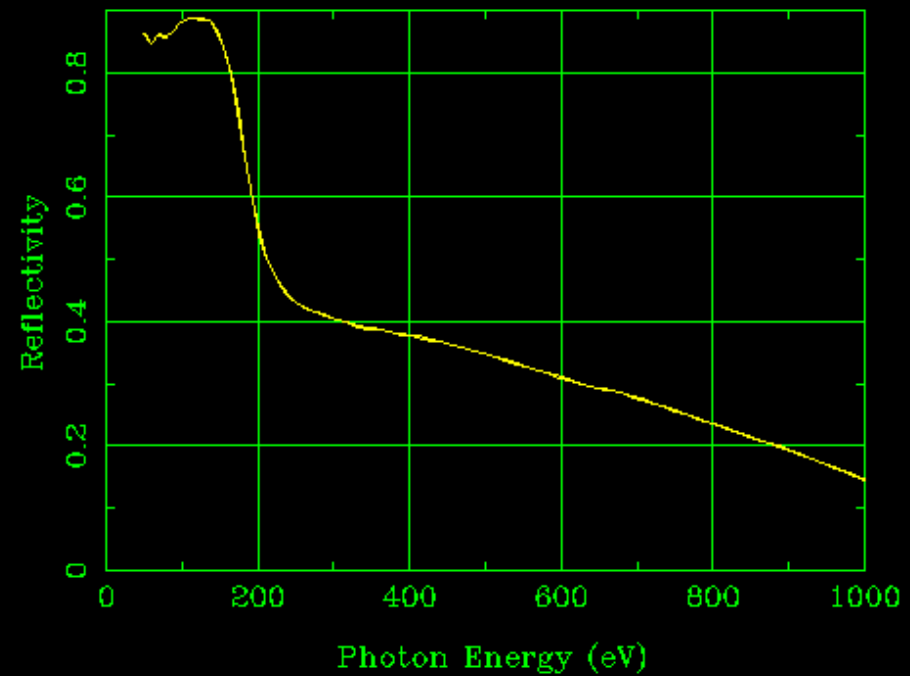
Protection against oxidation etc.

Typical coating material in VUV-SX: Au

88 deg



86 deg

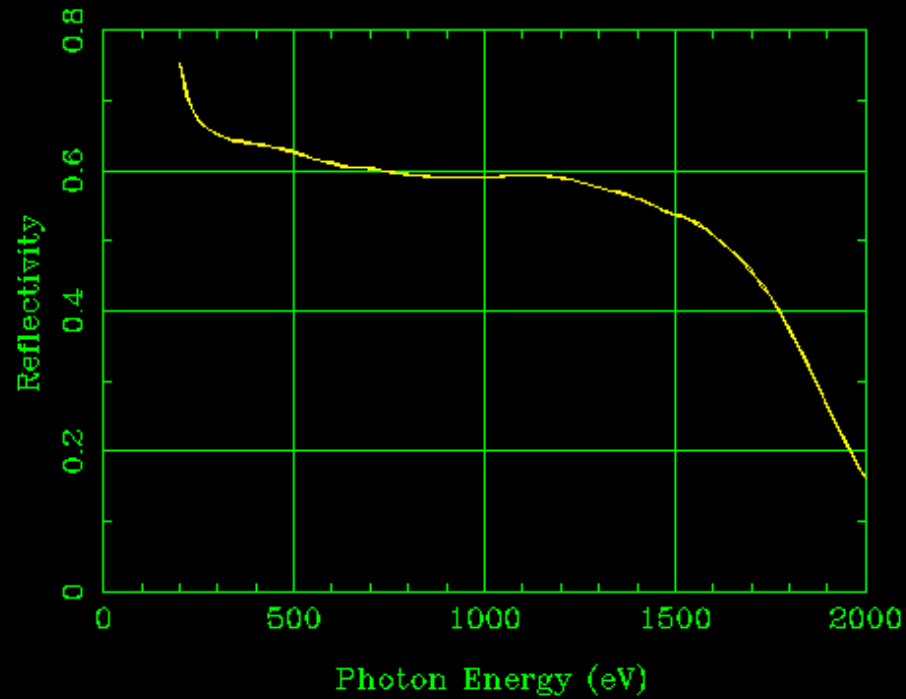


[http://henke.lbl.gov/optical\\_constants/](http://henke.lbl.gov/optical_constants/)

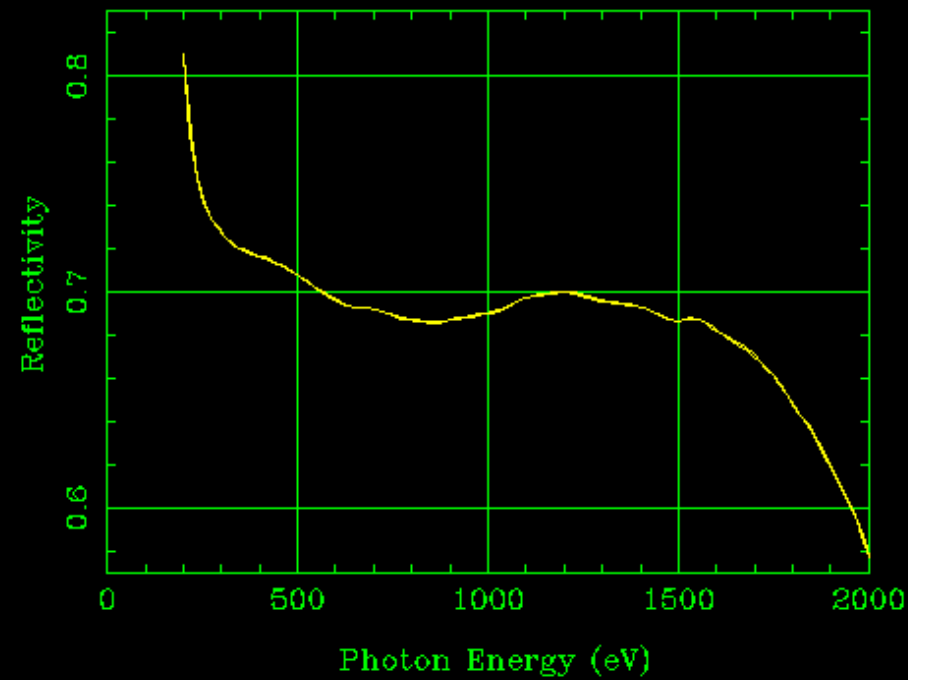
## 1.3. coating & reflectivity

Higher energy region

Au: 88 deg



Au: 88.5 deg

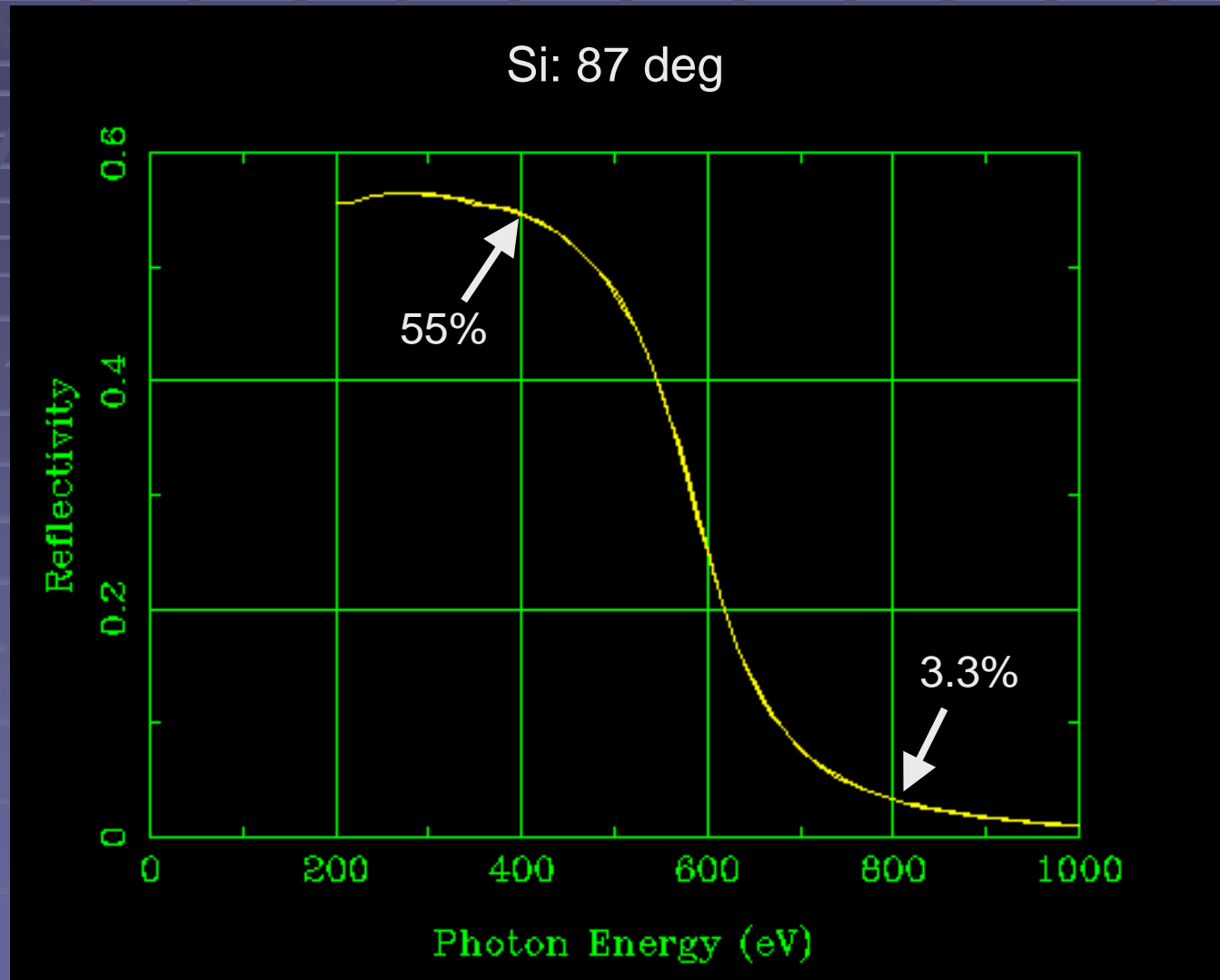


[http://henke.lbl.gov/optical\\_constants/](http://henke.lbl.gov/optical_constants/)



### 1.3. coating & reflectivity

Higher order suppression

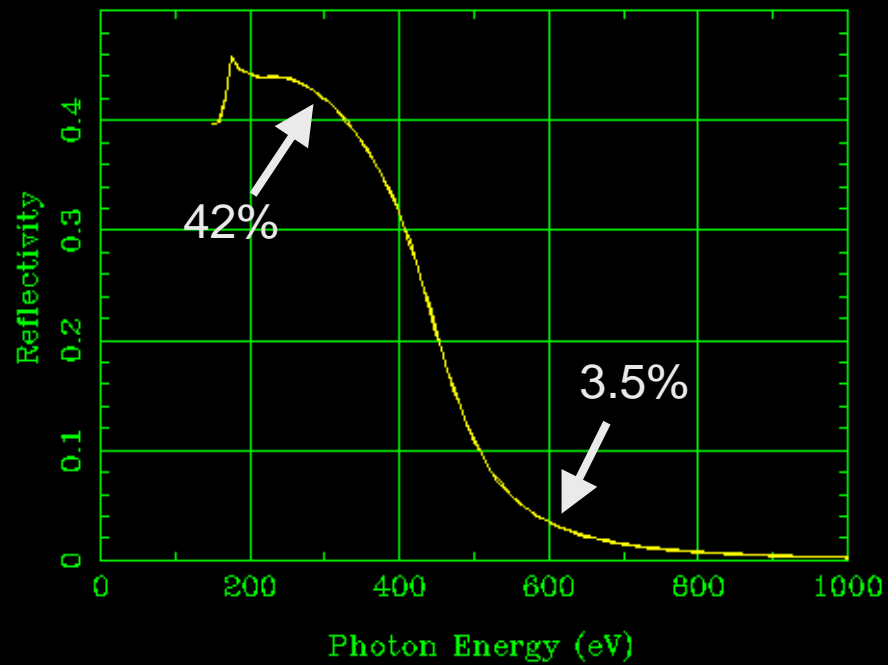


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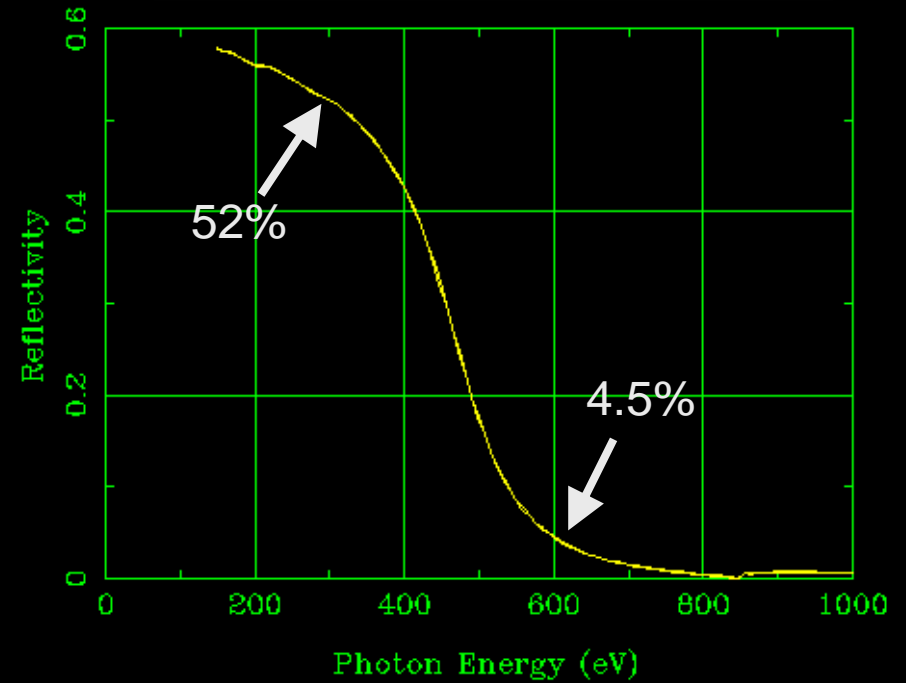
## 1.3. coating & reflectivity

### Higher order suppression

Si: 86 deg



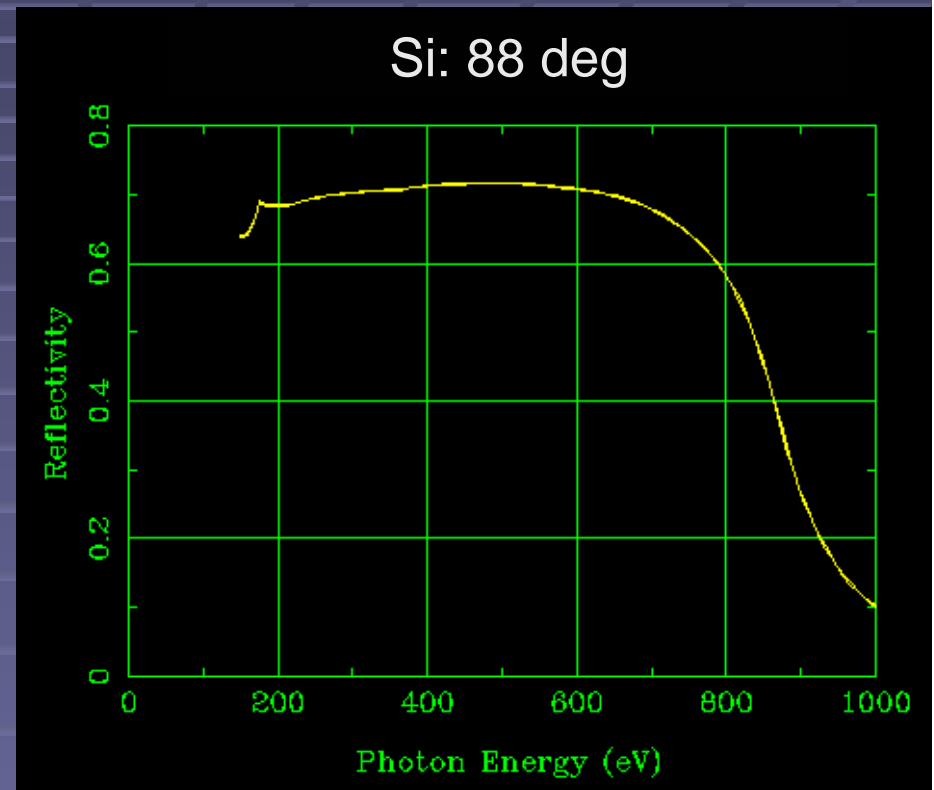
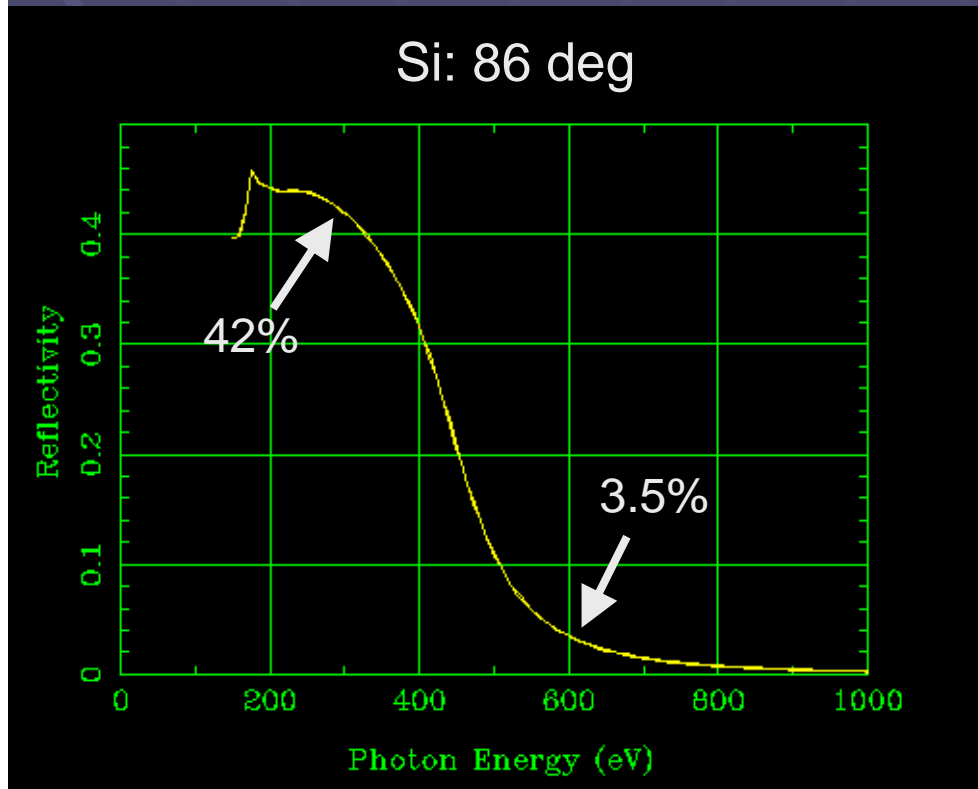
Ni: 84 deg



[http://henke.lbl.gov/optical\\_constants/](http://henke.lbl.gov/optical_constants/)

### 1.3. coating & reflectivity

Higher order suppression: **incidence angle dependence**

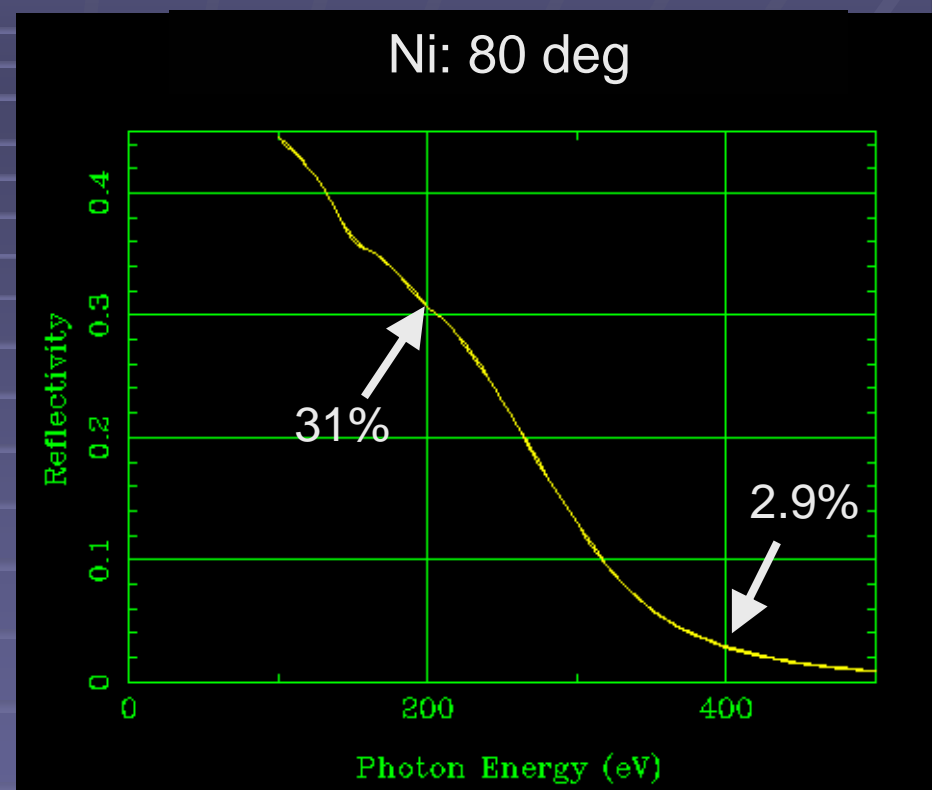
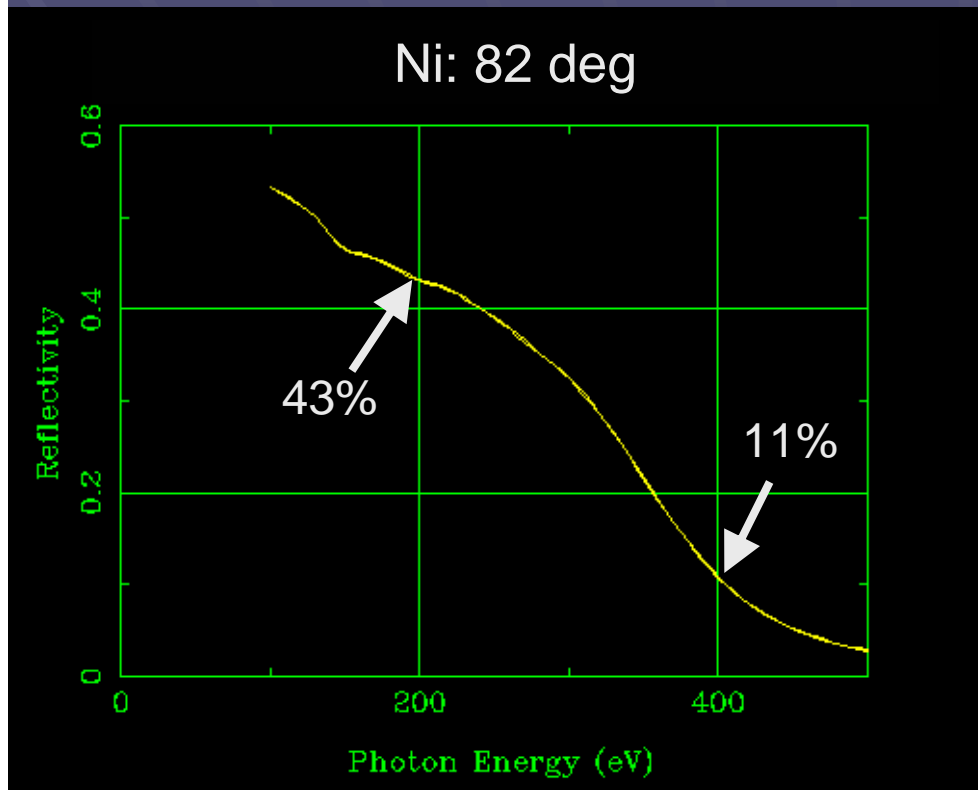


[http://henke.lbl.gov/optical\\_constants/](http://henke.lbl.gov/optical_constants/)

**Simultaneous rotation** with energy scan is necessary  
**No precise control** is required

### 1.3. coating & reflectivity

Higher order suppression in **low energy region**



[http://henke.lbl.gov/optical\\_constants/](http://henke.lbl.gov/optical_constants/)

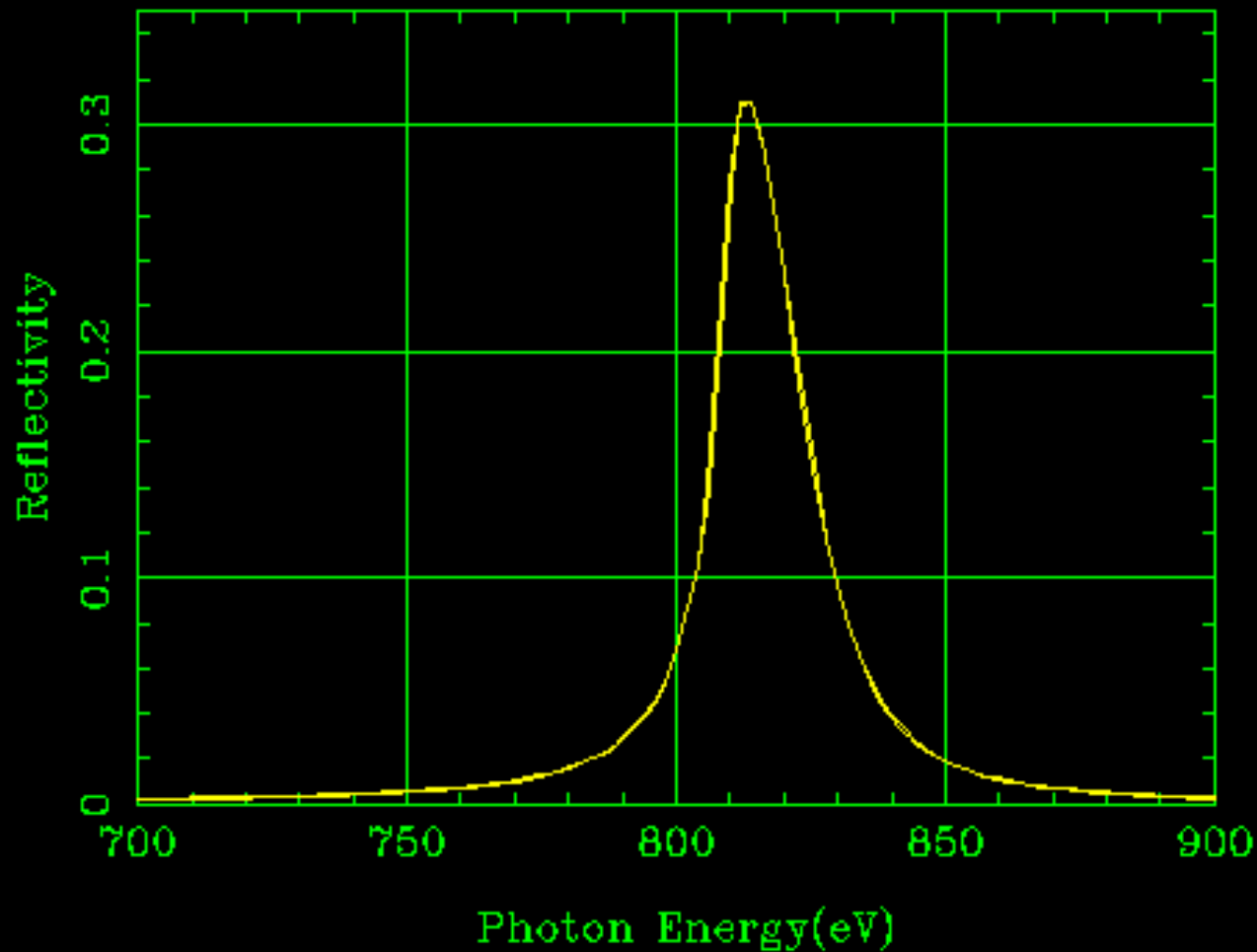
It is difficult to achieve high reduction ratio  
keeping high reflectivity for fundamental light

**Not effective below ~100 eV**

### 1.3. coating & reflectivity

#### Multilayer mirror

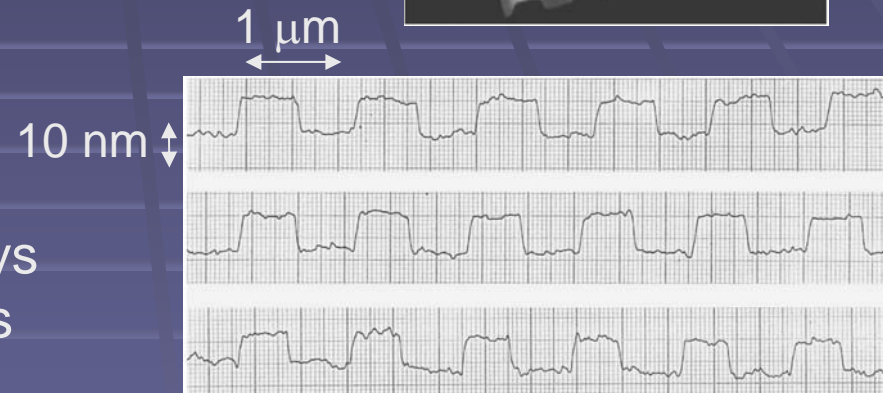
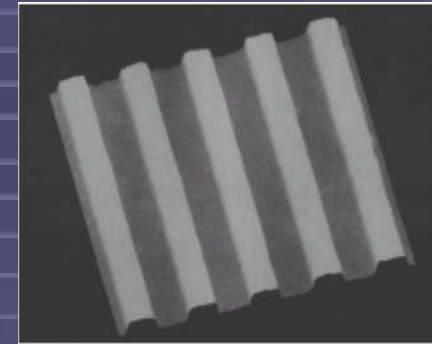
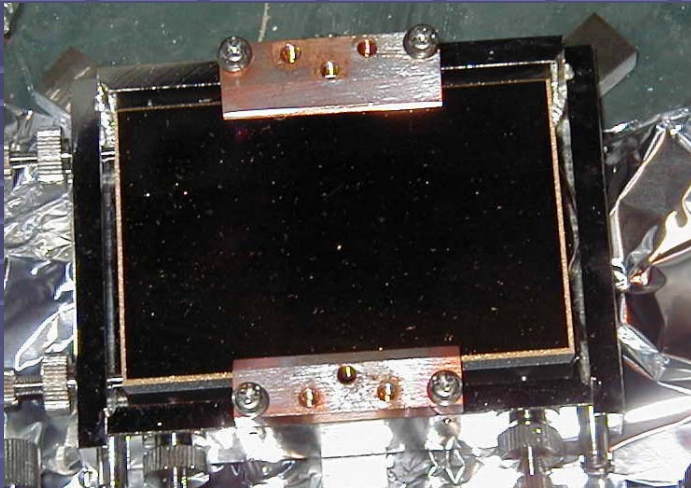
$[W/C]_{80}$ ,  $d = 3$  nm,  $75$  deg



[http://henke.lbl.gov/optical\\_constants/](http://henke.lbl.gov/optical_constants/)

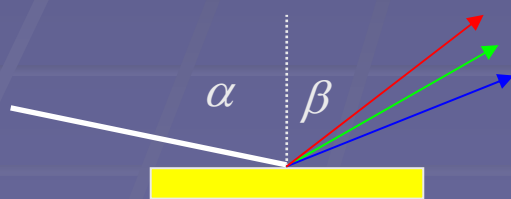
## 2.1. principle - wavelength dispersion -

Diffraction grating: **Periodic grooves** on a substrate



Principle: Interference between the rays reflected at different grooves

Enhanced when **light path difference =  $m\lambda$**



\*  $\alpha > 0$

$$\sin \alpha + \sin \beta = nm\lambda$$

$$* \alpha \neq |\beta|$$

(if  $\alpha = |\beta|$ , any  $\lambda$  satisfies the above condition at  $m = 0$ )

$n$ : groove density

$m$ : diffraction order

⇒ **zero-th order light**

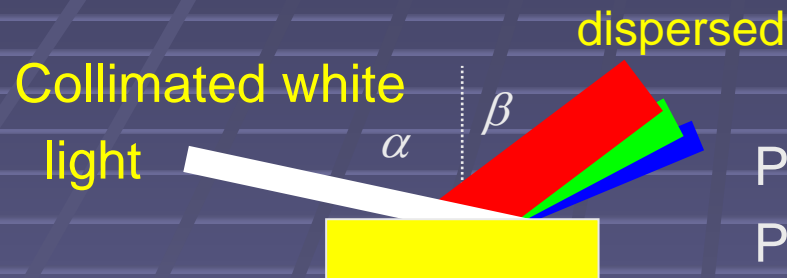
$\beta$  depends on  $\lambda$  ⇒ **Wavelength dispersion**

(conversion of wavelength to angle)

## 2.2. energy resolution - dispersion and focus -

How can we **monochromatize** by using a **diffraction grating**?

Most **basic** mode: **collimated-light illumination**



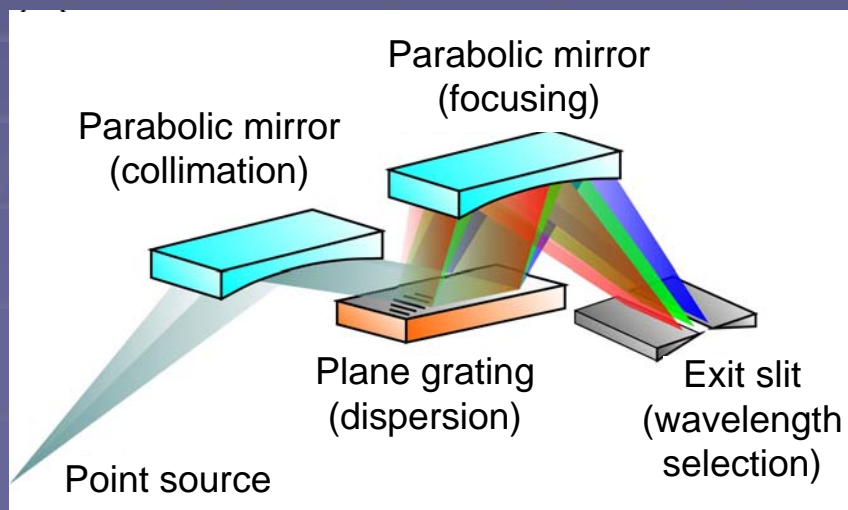
$$\sin \alpha + \sin \beta = nm\lambda$$

Problem 1: SR is **not a collimated light** !

Problem 2: **Superposition** of diffracted lights  
 $\Rightarrow$  **difficult to be resolved**

Solution 1: **Collimation** of diverging light with a **parabolic mirror**

Solution 2: **Focusing** of diffracted lights with another **parabolic mirror**

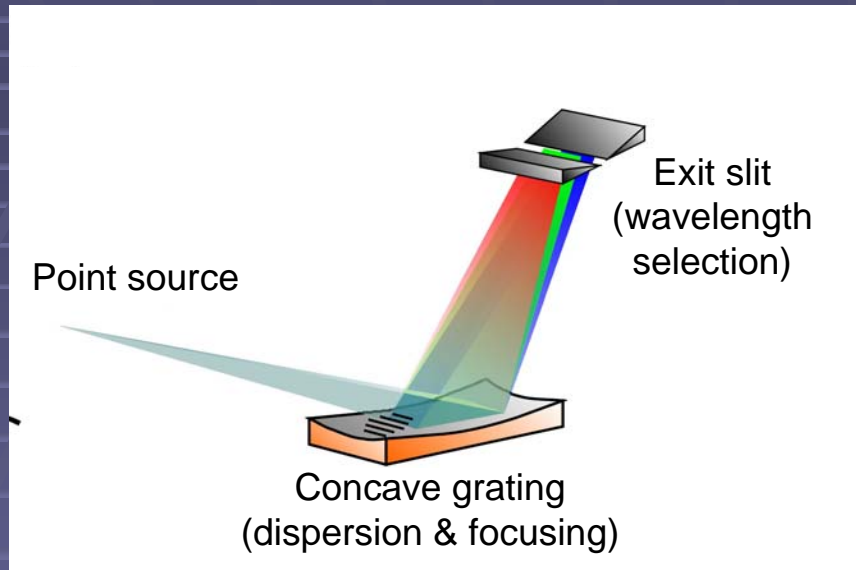


**Focused** diffracted lights are **well resolved** in wavelength at the exit slit !

**Dispersion and Focus**

## 2.2. energy resolution - dispersion and focus -

The **simplest** monochromator



Both the “**dispersion**” and “**focus**” are achieved by a **diffraction grating** only.

Is that really possible?

It is impossible to obtain a perfect focus at all wavelength

But, “perfect focus” is not necessary !

Small number of optical elements



## 2.2. energy resolution - dispersion and focus -

What determines the energy resolution?

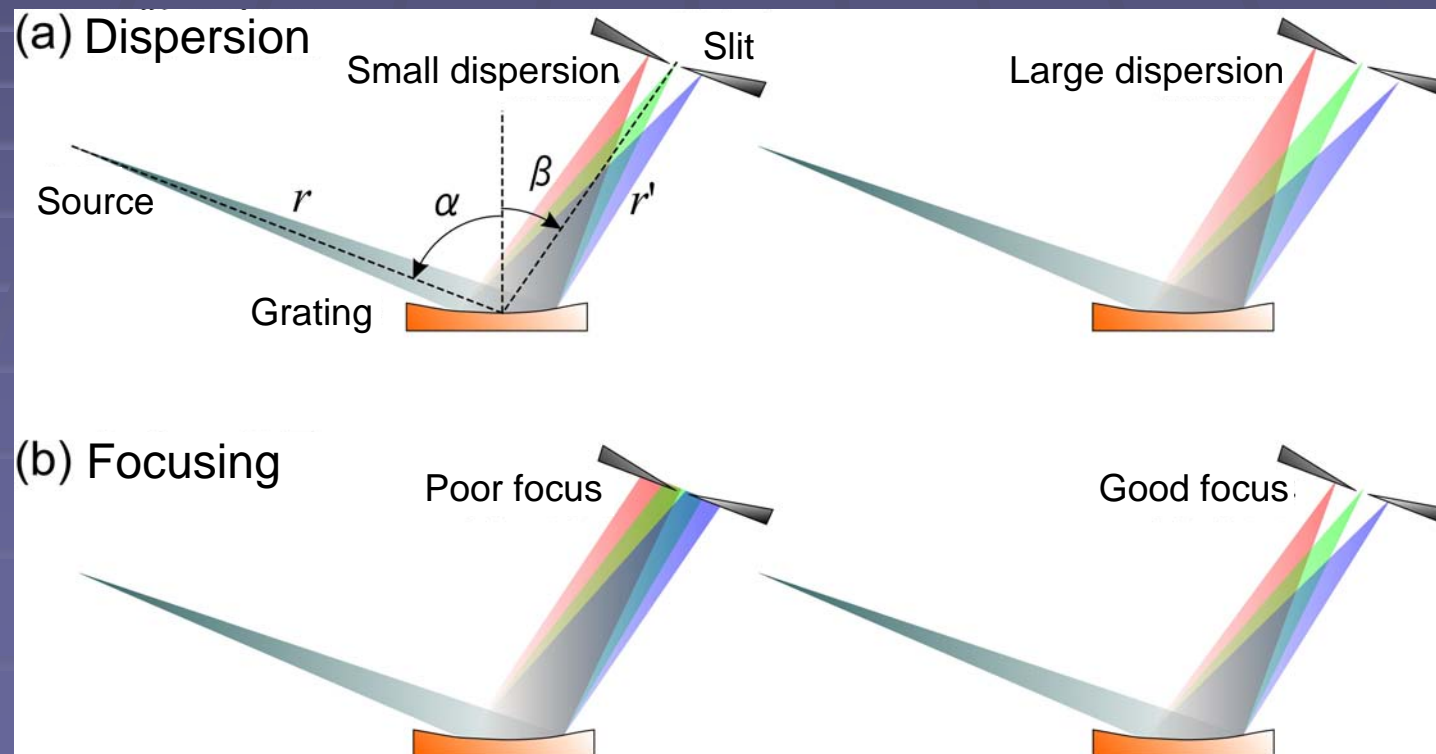
“Dispersion”: separation of lights with different wavelengths  $\Rightarrow dz/d\lambda$

$$\sin \alpha + \sin \beta = nm\lambda \Rightarrow$$

$$\cos \beta d\beta = nmd\lambda, \quad dz/d\lambda = r' d\beta / d\lambda = r' nm / \cos \beta$$

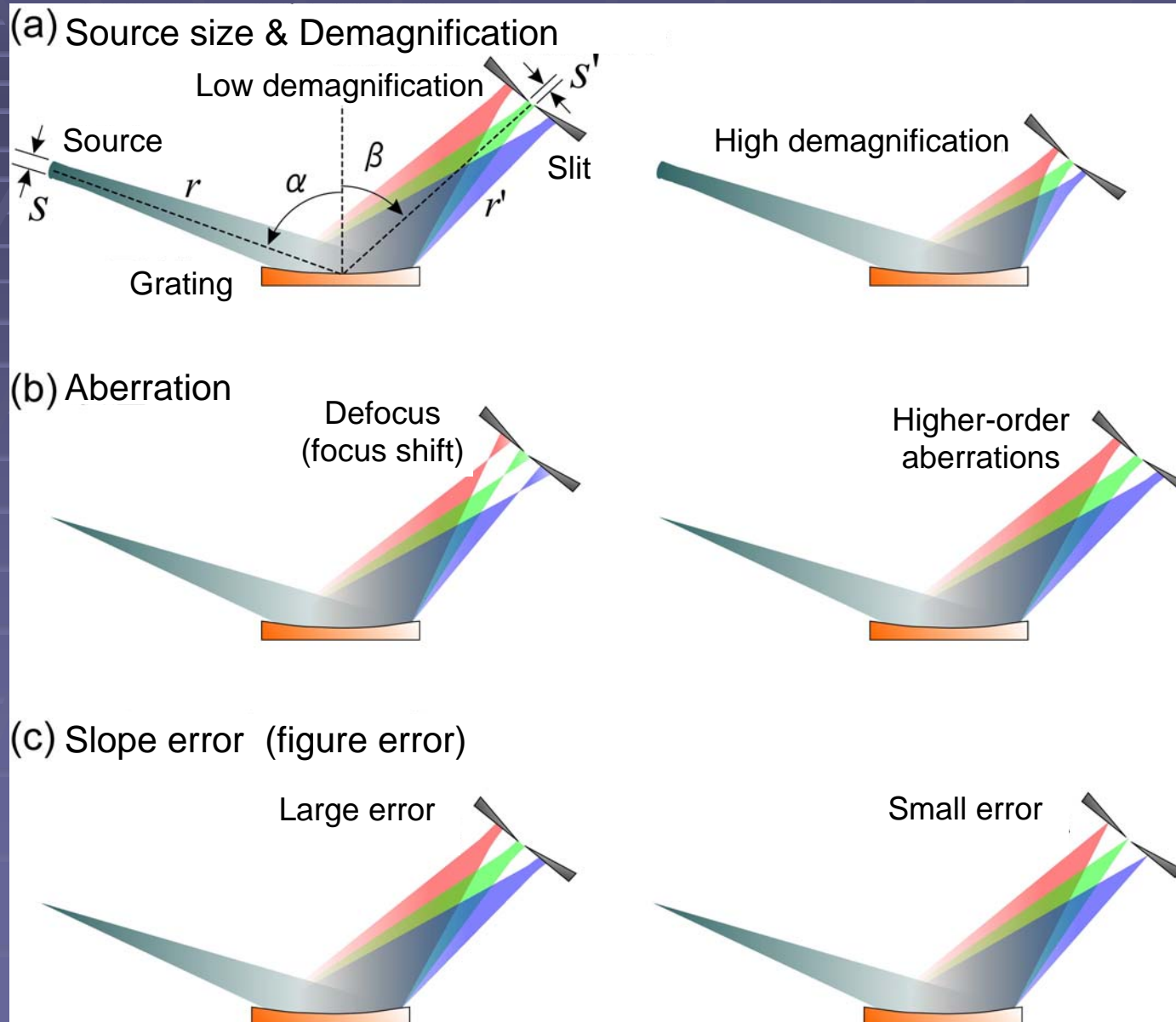
Ratio between “dispersion” and “light size” determines resolution.

i.e. large dispersion & small focus  $\Rightarrow$  high energy resolution



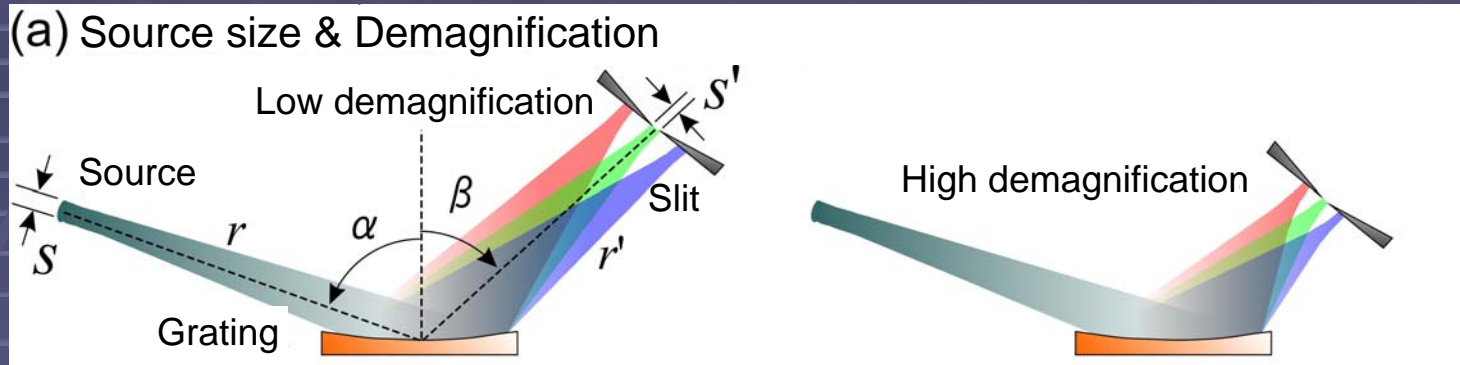
## 2.2. energy resolution - dispersion and focus -

What determines the **focus size**?



## 2.2. energy resolution - dispersion and focus -

### (a) Source size & Demagnification



$$s' \text{ (lower limit)} = s d/d'$$

$d$ : **divergence** at the source

$d'$ : **divergence** at the focus

How to **reduce**  $s'$ :

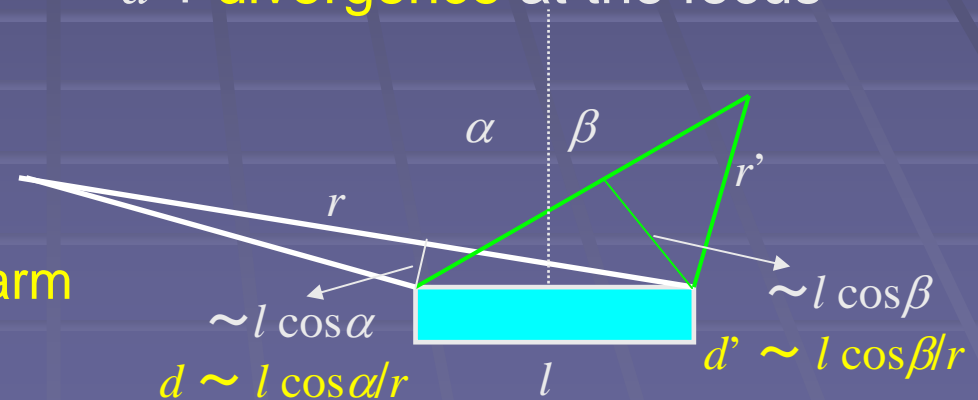
1. reduce  $r'$  compared to  $r$

**long entrance arm & short exit arm**

2. decrease  $|\beta|$  compared to  $\alpha$

**make the incidence angle more grazing**

(keeping the diffraction condition satisfied)



$\Rightarrow$  **High groove density**

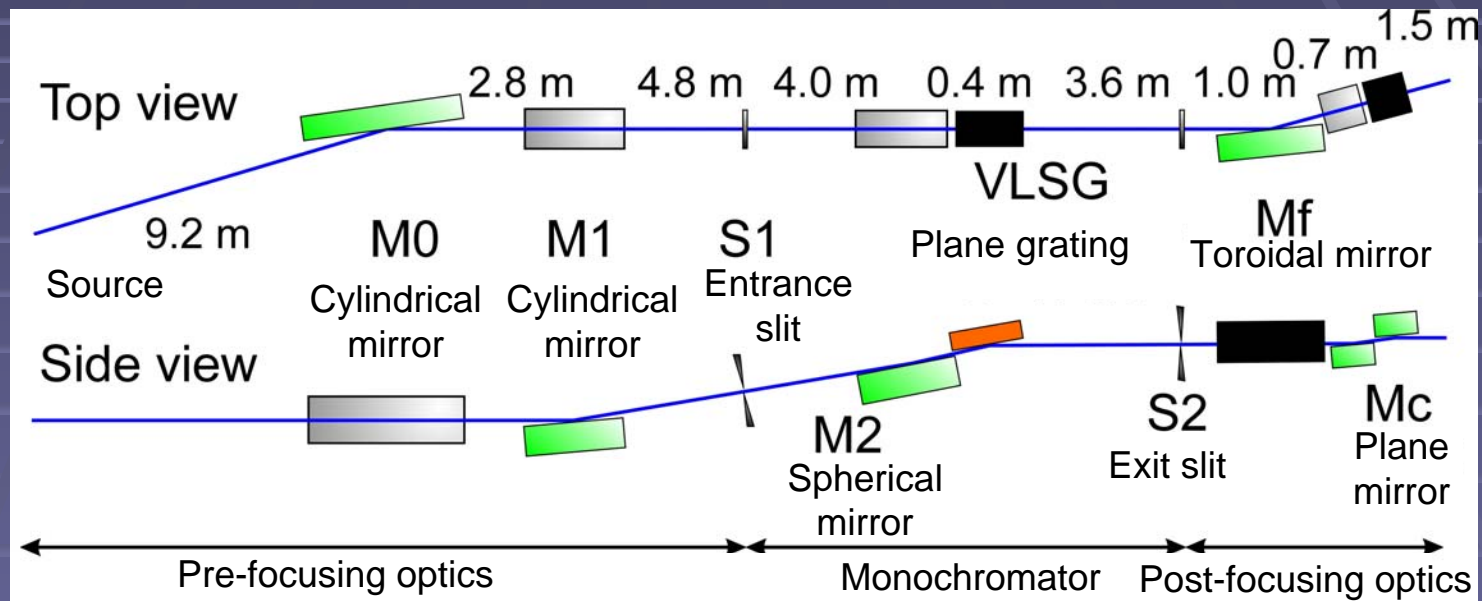
**large included angle** ( $\alpha - \beta$ )

## 2.2. energy resolution - dispersion and focus -

### 3. reduce the source size ( $s$ )

Must we reduce the size of the SR source itself?

Use of “pre-focusing optics” and “entrance slit”



“entrance slit” can be regarded as a **virtual source**

Source size can be controlled by entrance-slit opening  
(at the sacrifice of intensity)

Demagnification in the pre-focusing optics is effective  
(\* divergence increases)

## 2.2. energy resolution - dispersion and focus -

### (b) Aberration

Caused by a deviation from the elliptical (or parabolic) shape

e.g. Use of a spherical mirror instead of elliptical one

**Diffraction effects** should be taken into account for a diffraction grating

Aberration is usually expanded in a power series of the position on the optical element, ( $w, l$ ), using the **light path function,  $F$**

$$\begin{aligned}
 F &= p_A + q_A + r_B \\
 &+ M_{10}w + (M_{20}w^2 + M_{02}l^2 + M_{30}w^3 + M_{12}wl^2)/2 \\
 &+ (M_{40}w^4 + M_{22}w^2l^2 + M_{04}l^4)/8 + \dots \\
 &+ [n_{10}w + n_{20}w^2/2 + n_{30}w^3/2 + (n_{40}w^4)/8 + \dots] m\lambda, \quad \leftarrow \text{For grating}
 \end{aligned}$$

Diffraction condition

Defocus

(deviation from focal condition)

“coma” aberration

$$M_{10} = -\sin \alpha - \sin \beta,$$

$$M_{20} = (\cos^2 \alpha)/r_A + (\cos^2 \beta)/r_B,$$

$$M_{30} = (\sin \alpha \cos^2 \alpha)/r_A^2 + (\sin \beta \cos^2 \beta)/r_B^2$$

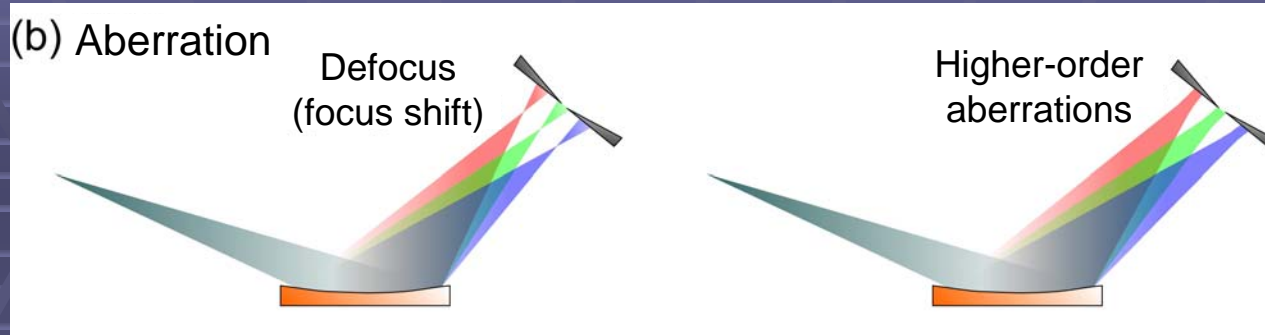
(sagittal focusing) or

$$M_{30} = (\sin \alpha \cos^2 \alpha)/r_A^2 + (\sin \beta \cos^2 \beta)/r_B^2 - [2(A_{10})_A^2 K_A]/R_A$$

(meridional focusing),

**Larger illumination area (larger  $w$  and  $l$ )  $\Rightarrow$  larger effects of aberration**

## 2.2. energy resolution - dispersion and focus -



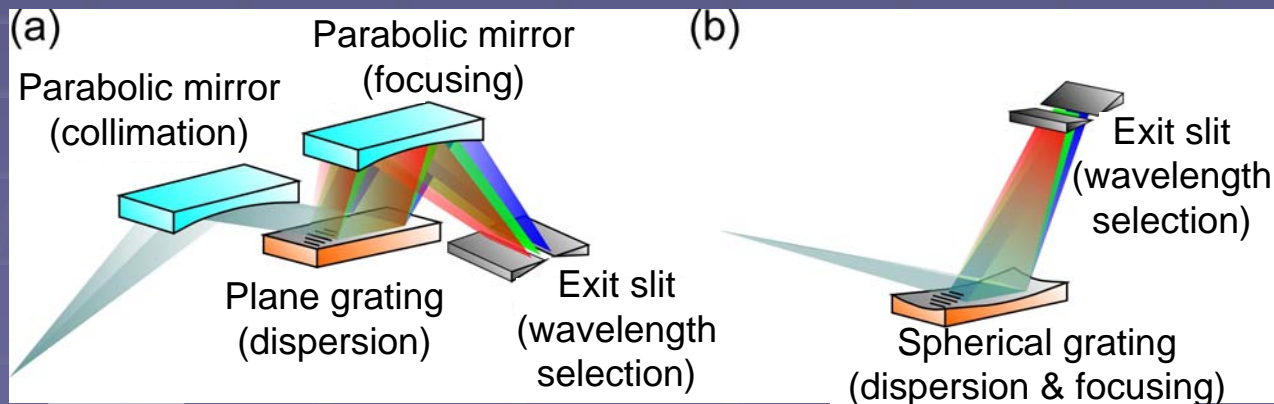
How to reduce the aberration:

**Defocus** can be **compensated** by adjusting the exit-slit position

Higher-order aberrations can be canceled by a combination of mirrors (not easy)

Reduce the illumination area  $\Rightarrow$  small divergence (acceptance)

Examples for aberration-free or low-aberration optics



**No aberration**

**“Rowland condition”**

$$r = R \cos \alpha, r' = R \sin \beta$$

## 2.2. energy resolution - dispersion and focus -

### (c) Slope errors

**Errors in fabrication** of optical elements (e.g. undulation of a plane mirror)

Not systematic  $\Rightarrow$  compensation is impossible

$\Rightarrow$  One have to fabricate optical elements with small slope error

or design optics to reduce the effects of the slope errors

**Demagnification** is also effective to reduce the slope-error effects

### (d) Number of illuminated grooves

Intrinsic problem of diffraction

**Resolving power**  $(\lambda/\Delta\lambda) \sim N$

\* Small divergence

$\Rightarrow$  small effects of aberration but small number of grooves

## 2.3. fabrication of gratings

1. **Substrate** (Plane, Cylindrical, Spherical, Toroidal,...)

Same as mirror fabrication

2. **Fabrication of grooves** (uniform or varied line spacing)

$$N = N_0 (1 + a_1 w + a_2 w^2 + a_3 w^3) \quad (\text{groove density})$$

$$\begin{aligned} F = & p_A + q_A + r_B \\ & + M_{10}w + (M_{20}w^2 + M_{02}l^2 + M_{30}w^3 + M_{12}wl^2)/2 \\ & + (M_{40}w^4 + M_{22}w^2l^2 + M_{04}l^4)/8 + \dots \\ & + [n_{10}w + (n_{20}w^2)/2 + (n_{30}w^3)/2 + (n_{40}w^4)/8 + \dots] m\lambda, \end{aligned}$$

= 0 for uniform line spacing

- (a) **Mechanical ruling**

All groove parameters ( $a_1, a_2, \dots$ ) can be **controlled** !

Relatively **rough surface**  $\Rightarrow$  causes **stray light**

Suitable for “**Brazed**” gratings



## 2.3. fabrication of gratings

### (b) Holographic recording

Interference patterns of Laser lights

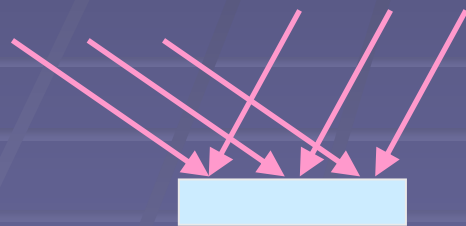
Some groove parameters might not be **controlled**

\* aspheric wavefront recording is available

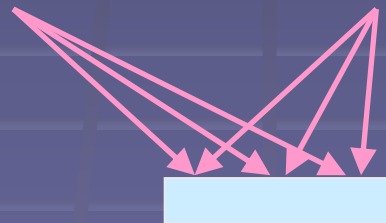
Relatively **smooth surface**  $\Rightarrow$  **high reflectivity** & **low stray light**

Both the “**Laminar**” and “**Brazed**” gratings can be fabricated

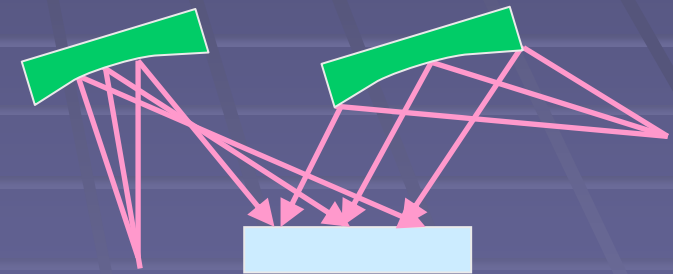
\* some manufacturer strongly prefers the Laminar type



Collimated lights  
 $\Rightarrow$  **uniform line spacing**



Spherical wavefronts  
 $\Rightarrow$  **varied line spacing**  
(poor control)



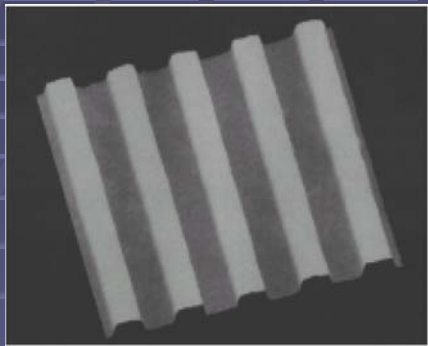
Aspheric wavefronts  
 $\Rightarrow$  **varied line spacing**  
(fine control)

T.Namikoka and M.Koike  
Appl. Opt. 34 (1995) 2180

## 2.3. fabrication of gratings

### 3. Groove shape (Laminar & Brazed)

#### (a) Laminar type

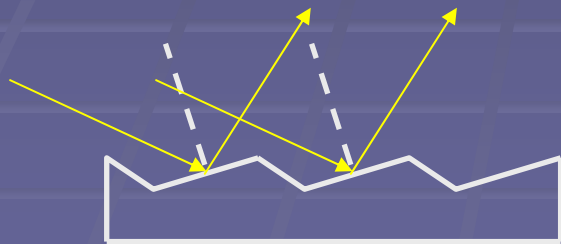


Medium diffraction efficiency

Higher order suppression

interference between top and bottom parts

#### (b) Brazed type



High diffraction efficiency when “on Braze”

Strong higher orders